ERRATUM: "THE 3.3 μm PAH FEATURE AS A SFR INDICATOR: PROBING THE INTERPLAY BETWEEN SF AND AGN ACTIVITIES" (PKAS, 27, 281[2012])

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Section 3 of the article "The 3.3 μm PAH Feature as a SFR Indictor: Probing the Interplay between SF and AGN Activities" which appears in the Volume 27, Number 4, pp 281~284, summarizes Woo et al. (2012). Therefore, we made the following modifications to the original article.

1. Page 3, section 3

The section title is modified, and Footnote 1 is added at the bottom of the page as the following.

3. A Case Study of z \sim 0.36 Seyfert $1s^1$

¹ This section summarizes the findings of Woo et al. (2012).

2. Page 3

Figures 2 and 3 are modified along with their captions as the following on the right column.

 $3. \ {\rm Page} \ 4$

A reference is added:

Woo, J. -H., Kim, J. H., Imanishi, M., & Park, D., 2012, AJ, 143, 49

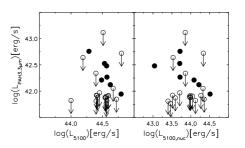


Figure 2: The correlations of $L_{PAH3.3}$ with the restframe 5100 Å luminosity $(L_{5100})(left)$ and nuclear AGN luminosity based on $L_{5100}(right)$. While filled circles represent the PAH detected galaxies, open circles with downward arrows represent upper limits of $L_{PAH3.3}$ for the undetected galaxies (a reproduction of Figure 3 of Woo et al., 2012).

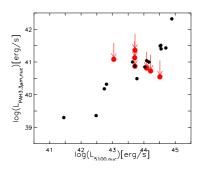


Figure 3: The correlation between the nuclear 3.3 μm PAH luminosity (L_{PAH3.3,nuc}) and AGN optical luminosity. Red filled circles represent L_{PAH3.3,nuc} estimated from L_{PAH3.3} of the AKARI sample, while black filled circles represent L_{PAH3.3} measured from ground-based long-slit spectroscopy (a reproduction of Figure 5 of Woo et al., 2012).