# Properties of Photoluminescience for AgInS<sub>2</sub>/GaAs Epilayer Grown by Hot Wall Epitaxy

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The AgInS $_2$  epilayers with chalcopyrite structure grown by using a hot-wall epitaxy (HWE) method have been confirmed to be a high quality crystal. From the optical absorption measurement, the temperature dependence of the energy band gap on the AgInS $_2$ /GaAs was derived as the Varshni's relation of  $E_g(T) = 2.1365$  eV -  $(9.89 \times 10^{-3} \text{ eV/K})$   $T^2$ /(2930 +T eV). After the as-grown AgInS $_2$ /GaAs was annealed in Ag-, S-, and In-atmosphere, the origin of point defects of the AgInS $_2$ /GaAs has been investigated by using the photoluminescence (PL) at 10 K. The native defects of  $V_{Ag}$ ,  $V_{S}$ ,  $Ag_{int}$ , and  $S_{int}$  obtained from PL measurement were classified to donors or acceptors type. And, we concluded that the heat-treatment in the S-atmosphere converted the AgInS $_2$ /GaAs to optical p-type. Also, we confirmed that the In in the AgInS $_2$ /GaAs did not form the native defects because the In in AgInS $_2$  did exist as the form of stable bonds.

Keywords: Hot-wall epitaxy, Optical absorption, Point defects, Annealing effects, Photoluminescience

## 1. INTRODUCTION

Ternary chalcopyrite crystals are currently of technological interest since they show promise for application in the areas of visible and infrared lightemitting diodes, infrared detectors, optical parametric oscillators, upconverters, and far infrared generator. But, many of these materials were reported to be difficult to grow in the form of a large and high quality crystal. Also, the fundamental physical properties of these compounds are still limited.

AgInS<sub>2</sub> is a ternary compound semiconductor, which has a wide band gap and belongs to the visible region of the spectrum. Its structure was crystallized in the form of chalcopyrite. Thereby, the AgInS<sub>2</sub> has uniaxial lattice structure[1]. Also, it can be made usefully n-type[2]. The AgInS<sub>2</sub> is one of the interesting materials used as optoelectronic devices applicable to the visible region. Many of the fundamental properties of the AgInS<sub>2</sub> such as photoconductivity[3], heat capacity[4], Hall effect[5], the calculation of the band structure[6], and optical absorption[7] have been carried out. Some attempts have also been made with the object of improving the efficiency of junction devices formed in the ternary compounds. But, for device applications of AgInS<sub>2</sub>, it is vital to know the electro-optical properties of this

material. The electronic and optical properties of  $AgInS_2$  crystals are mainly determined by point defects associated with individual atoms forming the ternary compound. However, optical properties and single layer growth of  $AgInS_2$  have not been well understood. Only, several researchers[8,9] have been investigated the electrical properties and polycrystalline  $AgInS_2$  film growth.

Sulfur is the higher vapor pressure compared to these of silver and indium. This is strongly related to native defects generated by non-stoichiometric composition during high temperature growth. These native defects, such as sulfur vacancy (V<sub>S</sub>), silver vacancy (V<sub>Ag</sub>), sulfur interstitial (S<sub>int</sub>), silver interstitial (Ag<sub>int</sub>) and complex of these point defects have known to be produced when the crystal cooled down after the crystal growth. Among the defects, the V<sub>S</sub> and Ag<sub>int</sub> are an plausible defects because they act as donors. Other defects such as VAg and S<sub>int</sub> may form deep levels and/or acceptors. Consequently, low-temperature crystal growth and thin film deposition methods have been recognized as the key technology to reduce native defects in AgInS2. HWE method is one of the low-temperature crystal growth. Although the AgInS<sub>2</sub> grown by HWE has not been reported in the literature, we tried to grow the AgInS<sub>2</sub> epilayers using HWE in this work. HWE has been

known to be especially designed to grow epilayers under the condition of a close thermodynamics equilibrium[10].

In this paper, to estimate the predominant point defects of the as-grown AgInS<sub>2</sub> through various heat-treatment, we carried out measurements of the optical absorption and the PL spectra. Based on these results, we will discuss the origin of native defects of the AgInS<sub>2</sub>.

### 2. EXPERIMENT

Prior to epilayer growth, the polycrystalline AgInS<sub>2</sub> used as a source material for AgInS2 growth was formed as fellows. The starting materials; Ag, In, and S, were the shot forms of 6N purity. After these materials were weighed in stoichiometric proportions, these were sealed in a quartz tube which coated the carbon. The sealed ampoule was placed in the synthesis furnace and successively rotated at rate of 1 revolution per minute. In order to avoid the explosion of the ampoule due to the sulfur vapor pressure, the temperature of ampoule was increased gradually to 1050 °C. This temperature was then maintained for 48 h. After the growth of the polycrystalline AgInS<sub>2</sub> ingot, AgInS<sub>2</sub> epilayers were grown on semi-insulating (100) GaAs by HWE method using the grown AgInS<sub>2</sub> ingot as source materials[11]. The grown AgInS<sub>2</sub>/GaAs were analyzed by the double crystal x-ray diffraction (Bede Scientific Co. FR 590) to obtain the optimum growth condition. The most suitable substrate and source temperature to grow the undoped AgInS₂/GaAs were 410 °C and 680 °C, respectively. The minimum value of a full width at half maximum (FWHM) of the grown AgInS<sub>2</sub> was 121 arcsec obtained from the x-ray rocking curves. The thickness of the asgrown AgInS<sub>2</sub>/GaAs measured by a α-step profilometer (Tenco,  $\alpha$ -step 200) was 2.6  $\mu$ m. And, the chalcopyrite structure was confirmed by x-ray diffraction study. From the Hall effect on the as-grown AgInS<sub>2</sub>/GaAs at 293 K measured by the van der Pauw method, we obtained 9.35  $\times$  10<sup>17</sup> cm<sup>-3</sup> and 294 cm<sup>2</sup>/V · s for carrier density and mobility, respectively, which confirmed the as-grown epilayer to be a n-type semiconductor. The optical absorption measurement was performed with a UV-VIS-NIR spectrophotometer (Hitachi, U-3501) for a range of 400 nm to 800 nm with the temperature varied from 10 K to 293 K. To prepare samples of AgInS<sub>2</sub>/GaAs:Ag (annealed in the Ag vapor atmosphere), AgInS2/GaAs:S (annealed in the S vapor atmosphere), AgInS<sub>2</sub>/GaAs:In (annealed in the In vapor atmosphere), the as-grown AgInS2/GaAs with each of Ag, S, and In shots were sealed in a quartz ampoule at  $\sim 10^{-6}$  torr. The samples of AgInS<sub>2</sub>/GaAs:Ag, AgInS<sub>2</sub>/GaAs:S, and AgInS<sub>2</sub>/GaAs:In were annealed for 1 h at 650 °C, for 30 min at 400 °C, and for 1 h 550 °C, respectively. The PL measurement at 10 K was carried out using a cryogenic helium refrigerator (AP, CSA-202B). The samples mounted on the cold finger of a cryostat were focused using the 442 nm line of He-Cd laser (Kimmon, 60mW). Then, the emitted light was detected by the photomultiplier tube through the monochromator. The detected signal was amplified by a lock-in amplifier and recorded in a x-y plotter.

### 3. RESULTS AND DISCUSSION

## 3.1 As-grown AgInS<sub>2</sub>/GaAs epilayer

Figure 1 shows the optical absorption spectra obtained in a temperature range between 10 K and 293 K. In order to identify the energy band gap for the AgInS<sub>2</sub>/GaAs, we carefully examined the relation between the optical absorption coefficient ( $\alpha$ ) and the incident photon energy (h  $\nu$ ) from the optical absorption measurements of the Fig. 1. The relation for a direct band gap between the h  $\nu$  and the  $\alpha$  is given as

$$(\alpha h \nu)^2 \sim (h \nu - E_g) \tag{1}$$

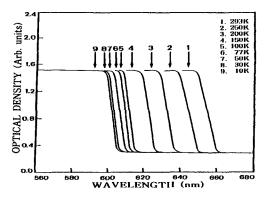


Fig. 1. Optical absorption spectra of the as-grown AgInS<sub>2</sub>/GaAs epilayer measured at different temperature.

Therefore, the temperature dependence of the optical energy band gap of our experimental, as shown in Fig. 2, is well described by the Varshni's equation[12]

$$E_g(T) = E_g(0) - \alpha T^2 / (T + \beta)$$
 (2)

where  $E_g(0)$  is the optical energy gap at absolute zero,  $\alpha$  and  $\beta$  are constants. From these experimental measurement, the  $E_g(0)$ ,  $\alpha$ , and  $\beta$  are determined to be 2.1365 eV, 9.89  $\times$  10<sup>-3</sup> eV/K, and 2930 K, respectively.

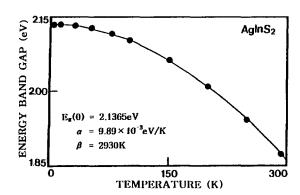


Fig. 2. Experimental values of the optical energy band gap as function of temperature in the range between 10 and 293 K.

Figure 3 shows the typical PL spectra of the as-grown AgInS<sub>2</sub>/GaAs measured at 10 K. From the expanded spectrum of Fig. 3, the two peaks at 593.8 nm (2.0880 eV) and 596.9 nm (2.0771 eV) appear on the shoulder toward the short-wavelength region. These peaks are free excitons associated with light-hole-exciton (lhx) and heavy-hole-exciton (hhx), respectively. But the lhx and the hhx may be considered as the upper polariton and the lower polariton of free exciton[13-16]. The splitting energy between the lhx and the hhx is 109 meV, which is caused by the strain due to the lattice mismatch between substrate and epilayer in the heteroepilayer growth. The equations of the hhx and the lhx are given by

$$hhx = E_{\varrho}(10) + \delta E x^{hh} \tag{3}$$

and

$$lhx = E_o(10) + \delta E x^{lh} \tag{4}$$

where  $\delta Ex^{hh}$  and  $\delta Ex^{lh}$  are the binding energy of the hhx and the lhx due to the variance of the strain at  $\vec{k}$  = 0, respectively. The value of  $E_g(10)$  was calculated to be 2.1362 eV using the eq. (2) and the binding energy of the free exciton,  $\delta Ex^{lh}$ , is obtained to be 48.2 meV. And the I<sub>x</sub> peak emitted due to the transition between free exciton to bound exciton is shown at 597 nm eV). The very strong intensity corresponding to the neutral donor bound exciton, I2, appears at 597.8 nm (2.0741 eV). This exciton is known to be a recombination from bound exciton to neutral donor. Here, the binding energy[17], E<sub>D</sub>, of the donorimpurity can be calculated by

$$I_{2}(h \ \nu) = E_{d}(10) - \delta E_{x}^{J_{1}} - 0.15E_{D} \tag{5}$$

 $E_{\rm D}$  is determined to be 92.7 meV. This can be ascribed to the donor states of  $V_{\rm S}$  or  $Ag_{\rm int}$ , which are located at 92.7 meV below the edge of the conduction band. At the same

time, this value is nearly equal to the activation energy obtained by Hall effect measurement. A neutral acceptor bound exciton, I<sub>1</sub>, of the sharp intensity peak at 600.1 nm (2.0660 eV) and LO phonon replica at 604.8 nm (2.0500 eV) appear on the right region of the wavelength. In the PL measurement, the observation of the free and bound excitons indicates that the as-grown AgInS<sub>2</sub> epilayer is a high quality because the emission peak of the exciton can be only observed under the condition of a long-range Coulomb coupling between the electron and the hole. And the donor-acceptor pair (DAP) emission at 609 nm (2.0358 eV) and its LO phonon replica were observed. Such DAP emission is caused by an interaction between donors and shallow acceptors. The very strong intensity and broad peak at 732.4 nm (1.6928 eV) in the longerwavelength region can be attributed to a self-activated (SA) emission.

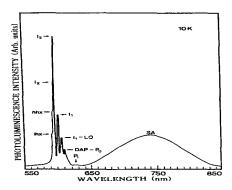


Fig. 3. Photoluminescence spectrum at 10 K for the asgrown AgInS<sub>2</sub>/GaAs epilayer.

## 3.2 Annealing effect of the AgInS<sub>2</sub>/GaAs epilayer

In order to know the origins of the several peaks of the as-grown AgInS<sub>2</sub>/GaAs, we measured the PL spectra for samples annealed in Ag-, S-, and In-atmosphere. Firstly, to know a role of Ag, we prepared AgInS<sub>2</sub>/ GaAs: Ag sample which were annealed in Ag-atmosphere for 1 h at 650 °C. Therefore, it would be saturated the AgInS<sub>2</sub>/GaAs with Ag to anneal the sample in Agatmosphere. The PL spectrum of AgInS<sub>2</sub>/GaAs:Ag measured at 10 K is shown in Fig. 4. By comparing the PL spectra of the as-grown AgInS2/GaAs as shown in Fig. 3 with these of Fig. 4, we found that the peaks related to the I<sub>1</sub> and its LO phonon replica completely disappeared in the AgInS<sub>2</sub>/GaAs:Ag. This disappearance indicates that the I<sub>1</sub> and its LO phonon replica are certainly associated with an acceptor level of VAg or levels of antisite native defects such as In<sub>Ag</sub> and Ag<sub>In</sub>. On the other hand, the I<sub>2</sub> emission became the dominant peak in the PL spectrum of the AgInS<sub>2</sub>/GaAs:Ag. This peak is known to be observed in the typical n-type AgInS<sub>2</sub>/GaAs. The FWHM value of the I<sub>2</sub> peak is taken to be 8 meV. However, the  $I_2$  peak is not related to Ag

because the sites of  $V_{Ag}$  should be substituted with the saturated Ag. Therefore, the  $I_2$  may be related to the  $V_S$  or  $Ag_{int}$ . Also, the SA emission completely disappeared in the sample, this means that the AgInS<sub>2</sub>/GaAs:Ag is refined by annealing in Ag-atmosphere and the SA emission is related to  $V_{Ag}$ . Among the samples prepared in this experiment, the DAP emission and its LO phonon replicas was dominantly observed only in the sample which was annealed in the Ag-atmosphere. The origin of these DAP emission may be associated with point defects of  $V_S$ ,  $Ag_{int}$ , or these complex. Also, the peaks of hhx, lhx, and Ix completely disappeared in the AgInS<sub>2</sub>/GaAs:Ag.

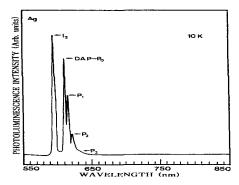


Fig. 4. Photoluminescence spectrum at 10 K for the AgInS<sub>2</sub>/GaAs:Ag epilayer.

Figure 5 displays the PL spectrum at 10K of the AgInS<sub>2</sub>/GaAs:S which was annealed in the S-atmosphere for 30 min at 400  $^{\circ}$ C. This figure shows that the I<sub>1</sub> peak is a very sharp and its PL intensity is high. However, the hhx, lhx, Ix, and I<sub>2</sub> are not observed. As is well known, small deviations from stoichiometry in the crystal can be adjusted by changes of the concentration of vacancies in the appropriate sublattice. Probably  $V_{Ag}$  and  $V_{In}$  are present in the AgInS<sub>2</sub>/GaAs which include excess S atoms. But, in the starting material to form AgInS2 epilayer, the concentration of In atoms is equal to that of Ag atoms. However, the concentration of V<sub>in</sub> will be little compared with that of  $V_{Ag}$ . Because the group I atoms participate only weakly in the covalent bonding and most of the covalent bonding occurred between the group III and group VI atoms[2]. Therefore, VAg are more likely than  $V_{ln}$  and  $S_{int}$ . The origin of  $I_1$  can be ascribed to the deep acceptor level originated from the V<sub>Ag</sub> due to stoichiometric deviation. And it means that the conversion from n-type to optical p-type occurs when  $V_S$  disappear. From results of the  $I_1$  emission, the binding energy of acceptor-impurity[17], E<sub>A</sub>, is obtained by

$$I_1(h \ \nu) = E_a(10) - \delta E x^{l_1} - 0.08 E_A$$
 (6)

E<sub>A</sub> is calculated to be 275 meV. The optical transitions of the E<sub>A</sub> may be associated with a deep level of S<sub>int</sub> or V<sub>Ag</sub> point defects. The intensity of the I<sub>2</sub> peak enhanced after the Ag-atmosphere treatment, however, it disappeared completely after the S-atmosphere treatment. Therefore, the origin of the I<sub>2</sub> emission is believed to be related to V<sub>S</sub>. The I<sub>1</sub> peak disappeared in the Agatmosphere treatment, it became the dominant peak after the S-atmosphere treatment. This indicates that the origin of the  $I_1$  emission is related to  $V_{Ag}$  or  $S_{int}$ . And the DAP emission and its LO phonon replicas were dominantly observed in the sample annealed in the S-atmosphere. Therefore, these peaks are related to  $V_{Ag}$ ,  $S_{int}$ , or these complex, which acted as acceptors. Also, the excess S atoms were contained in the AgInS2/GaAs:S, so that the V<sub>S</sub> were hardly contained in the AgInS<sub>2</sub>/GaAs:S. Ag atoms can be expected to readily move between interstitial positions, since they participate only weakly in the covalent bonding. Therefore, the most possible intrinsic donors are believed to be Agint. The Agint is related to the DAP emission, which acted as donors. Also, the intensity of the SA emission decreased after the S-atmosphere treatment. This means that the origin of the SA emission is related to  $V_S$ .

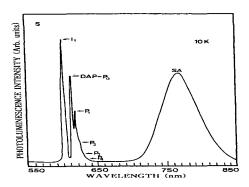


Fig. 5. Photoluminescence spectrum at 10 K for the AgInS<sub>2</sub>/GaAs:S epilayer.

Figure 6 shows the PL spectrum at 10 K of the AgInS<sub>2</sub>/GaAs:In annealed in the In-atmosphere for 1 h at 550 °C. With comparison of the peaks of Fig. 2, there is a similarity to respective peak positions and intensities of PL spectrum. This means that In is not related to form the native defect in the sample, because In forms more covalent bonds than Ag and S in chalcopyrite structure semiconductors, i.e., In is closer to the axis of the Periodic Table and participates in the formation of precovalent and less-ionic chalcopyrite compounds than Ag and S. Therefore, the above evidences indicate that In is the stable element in the AgInS<sub>2</sub>.

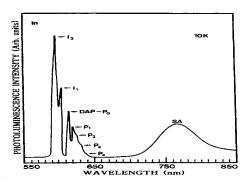


Fig. 6. Photoluminescence spectrum at 10 K for the AgInS<sub>2</sub>/GaAs: In epilayer.

### 4. CONCLUSION

The absorption and PL spectra of AgInS<sub>2</sub>/GaAs epilayers grown by using HWE method were investigated. The energy band gap obtained from the absorption spectra was well described by the Varshni's relation of  $E_g(T) = 2.1365 \text{ eV} - (9.89 \times 10^{-3} \text{ eV/K})$  $T^2/(2930 + T K)$ . The free excitons of the lhx and hhx have found in the as-grown AgInS<sub>2</sub>/GaAs and its splitting energy gap between the lhx and the hhx was determined to be 109 meV. This free exciton peak is associated with the strain due to lattice mismatch between substrate and epilayer. Also, the binding energy of the free exciton was estimated to be 48.2 meV. The I<sub>2</sub> emission was confirmed to be related to the V<sub>S</sub> or Ag<sub>int</sub> generated by non-stoichiometric composition. These defects were proved to be acted as donors. Therefore, these defects indicate one of the reasons why the AgInS<sub>2</sub> grown is generally the n-type. At the same time, the binding energy of the donor-impurity was calculated to be 92.7 meV.

The  $I_1$  emission became the dominant peak in the AgInS<sub>2</sub>/GaAs:S after the S-atmosphere treatment. We confirmed that the AgInS<sub>2</sub>/GaAs:S is converted into the optical p-type by using the PL measurement and that the origin of the  $I_1$  emission is related to  $V_{Ag}$  or  $S_{int}$ . The DAP emission is caused by interaction between donors such as  $V_S$  or  $Ag_{int}$ , and shallow acceptors such as  $V_{Ag}$  or  $S_{int}$ . Finally, the role of In in the AgInS<sub>2</sub>/GaAs is not related to forming the native defects because the In is the stable ingredient of the AgInS<sub>2</sub>.

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