# Quantum Mechanical Calculation of Spectroscopic Constants of CIO and CIO<sup>+</sup>

#### Hae-Sun Song and Eun-Mo Sung

Department of Science Education, Chungbuk National University, Chungbuk 306-763 Received February 22, 1993

The *ab* initio calculations were performed on ClO and ClO<sup>-</sup> using the configuration interaction and Møller-Plesset methods of several different levels of approximation. Three different basis sets, 66 contracted Gaussian-type orbitals, 6-31G<sup>\*</sup> and 6-311G<sup>\*</sup>, were employed in this calculation. The results of calculation were compared with the experimental values of ClO. The values from the calculation with 66cGTO basis set gave excellent agreement with the experimental values. The spectroscopic constants of ClO<sup>+</sup> were also predicted.

### Introduction

The role of chlorine monoxide molecule is thought to be a catalyst in the chlorine cycle for the destruction of ozone in the earth's stratosphere<sup>1</sup>. The importance of its role has resulted in much recent interest in CIO spectroscopically and theoretically. The emission spectra of CIO have been first observed by Pannetier and Gaydon<sup>2</sup> and the absorption spectrum by Poter<sup>3</sup>, Durie and Ramsay<sup>4</sup>, and Coxon Ramsay<sup>3</sup>. The molecular constants of CIO were assigned with the rotational analysis of the  $A^2H_i - X^2H_i$  band system by Coxon and Ramsay. Recently Andrews *et al.*<sup>6</sup> and Hudgens *et al.*<sup>7</sup> also performed the analysis of the absorption spectrum of CIO.

The rotational spectra of CIO radical in the millimeter and submillimeter wavelength region have been analyzed by Amano and Hirota<sup>8</sup>, and Kakar *et al.*<sup>9</sup> and the precise molecular parameters have been obtained. Lately the high-resolution infrared spectra of CIO radical have been observed using a Fourier transform spectrometer and a diode laser to calculate the fundamental transition moment of CIO(X<sup>3</sup>H<sub>i</sub>)<sup>10,11</sup>.

Because of the great interest about its catalytic role in the destruction of stratospheric ozone, numerous attempts have been made for the quantitative measurements of ClO. The techniques to measure atmospheric concentration of ClO have included balloon borne in situ method by atomic resonance fluorescence measurements<sup>12,13</sup>, millimeter wave heterodyne measurements of ClO emission lines<sup>14-16</sup> and infrared heterodyne measurements<sup>17-19</sup>.

There are a few theoretical calculations for ClO. O'Hare and Wahl<sup>20</sup> have calculated dissociation energy, ionization potential and dipole moment of ClO, ClO<sup>-</sup> and ClO<sup>-</sup>. Including the electron correlation energy, Lievin and Metz<sup>21</sup> have obtained  $R_e$ ,  $\omega_e$  and  $D_e$  of ClO. Langhoff *et al.*<sup>22</sup> have computed the electric dipole moment of ClO using the Cl method.

Although there have been a number of works for ClO, there is only one experimental work<sup>23</sup> for ClO<sup>+</sup>, which has observed the structure and vibrational frequencies of ClO<sup>+</sup>. Theoretical calculations of ClO<sup>+</sup> are also hardly found<sup>20</sup>.

These molecular ions and radicals are widely distributed in the interstellar medium and the information about those molecular ions have a great deal of effect to understand the chemistry and dynamics in the interstellar medium. And lately a lot of efforts have been concentrated to search those rare molecular ions<sup>24,25</sup>. The detection of interstellar molecules has been greatly aided by their spectroscopic constants from experiment or theory. Many of the molecular ions have been studied theoretically to obtain the spectroscopic properties and molecular geometries<sup>26–28</sup>.

In this work the molecular parameters of  $ClO^+$  have been calculated with M $\phi$ ller-Plesset(MP) perturbation method<sup>29</sup> and configuration interaction method(Cl), and the spectroscopic information has been provided for the further experimental search. The calculations on ClO radical were also performed to compare the data with known experimental values and to obtain the criteria for the error in calculations.

#### Calculation and Results

**CIO**<sup>+</sup>. All calculations were performed with the GAUS-SIAN 86 and 88 program. Electronic ground state of CIO is <sup>2</sup>*H* with an electronic configuration  $\cdots (6\sigma)^2 (7\sigma)^2 (2\pi)^4 (3\pi)^3$ . Therefore the electronic configuration of CIO<sup>+</sup> is likely to be  $\cdots (6\sigma)^2 (7\sigma)^2 (2\pi)^4 (3\pi)^3$ , and the lowest electronic state is <sup>3</sup> $\Sigma$ .

Several different levels of single point calculations were carried out. Basis sets of the 66cGTO (66 contracted Gaussian-type orbitals), 6-31G\* 30(single-zeta core, double-zeta valence, polarization functions on nonhydrogen atoms) and 6-311G\* (single-zeta core, triple-zeta valence, polarization functions on nonhydrogen atoms) were used in the calculations. The basis set of 66cGTO consisted of (10s, 6p, 2d) contracted to [7s, 4p, 2d] on the atom, O, and (13s, 9p, 2d) contracted to [9s, 6p, 2d] on the atom, Cl, where only the innermost functions of each symmetry were contracted. Huzinaga's primitives<sup>33</sup> were used for the (10s, 6p) set and the (12s, 9p) set of Veillard34 was modified for the (1s, 9p) set. The d exponents were related to the third most diffuse p exponents by successive multiplication by 1.5 and  $0.5^{35}$ . The (13s, 9p) set was obtained from the (12s, 9p) primitives of Veillard by replacing the two most diffuse s primitives with three by taking the geometric mean and multiplying by 2.5, 1.0 and 0.435. All electrons were correlated in the MP and Cl calculations.

In Figure 1 and 2 the total energies are shown for MP3, MP4SDQ and CISD. Figure 1 is obtained with the basis set of 6-311G\* and Figure 2 is obtained with the basis set of 66 cGTO. Using the basis set of the 66cGTO, the total energies were much lower than those of 6-31G\* and 6-311G\*



Figure 1. The potential curve of CIO<sup>+</sup> with 6-311G\* basis set.

for all levels of approximation.

The energy points were fit to the polynomials of sixth degree in the internal displacement coordinate  $\Delta R = R - R_c$ ;

$$V-V_0=\sum_{n=2}^6 f^n \Delta R^n/n !$$

The derived potential constants are listed in Table 1.

In addition, CI calculations were performed with the same basis sets. The CI calculation required about twice longer CPU time than that of MP calculation, but the energies of the CI calculations are higher than those of the MP methods. After the energy points were fit to the potential function,





Figure 2. The potential curve of CIO<sup>+</sup> with 66cGTO basis set.

the potential constants were obtained, and those are shown in Table 2.

From these potential function the various spectroscopic parameters were calculated by the method used in our previous paper<sup>31</sup>, Those values are in Table 3. The bigger basis set gave shorter bond length and smaller  $\omega$ , value. The CISD with 66cGTO gave short bond length, which is different from the previous results. The potential functions were different for various method of calculations and the spectroscopic constants were affected by the shape of potential function. For CIO<sup>+</sup> case no reference was available for the comparison purpose.

Since the molecular structure and spectroscopic properties

Table 1. The Potential Energy Function<sup>o</sup> of CIO<sup>+</sup> with the MP Energies

|                        | basis set | MP3        | MP4DQ       | MP4SDQ      |
|------------------------|-----------|------------|-------------|-------------|
| V <sub>0</sub> (a. u.) | 6-31G*    | -534.13759 | -534.13927  | - 534.14774 |
|                        | 6-311G*   | -534.19723 | - 534.19855 | -534.20823  |
|                        | 66cGTO    | -534.26483 | -534.26446  |             |
| R.(Å)                  | 6-31G*    | 1.47415    | 1.47382     | 1.48752     |
|                        | 6-311G*   | 1.46437    | 1.46396     | 1.47756     |
|                        | 66cGTO    | 1.46055    | 1.45722     | 1.46997     |
| $\int^2$               | 6-31G*    | 0.693906   | 0.658974    | 0.667586    |
|                        | 6-311G*   | 0.594924   | 0.591239    | 0.616537    |
|                        | 66cGTO    | 0.602948   | 0.554240    | 0.576580    |
| ſ³                     | 6-31G*    | -2.06669   | -2.10124    | - 1.76051   |
|                        | 6-311G*   | - 1.99502  | - 2.02123   | -1.74070    |
|                        | 66cGTO    | -1.91855   | -1.82761    | -1.61208    |
| $f^4$                  | 6-31G*    | 1.73136    | 1.81808     | 2.23093     |
|                        | 6-311G*   | 2.75814    | 2.82751     | 2.49935     |
|                        | 66cGTO    | 1.80786    | 2.82685     | 2.64915     |
| $f^{s}$                | 6-31G*    | - 18.3075  | - 17.4416   | - 25.1212   |
|                        | 6-311G*   | - 17.2389  | - 16.7670   | -21.7364    |
|                        | 66cGTO    | - 13.9487  | -17.4900    | -20.7169    |
| $f^6$                  | 6-31G*    | 157.228    | 154.342     | 140.911     |
|                        | 6-311G*   | 108.750    | 107.086     | 113.802     |
|                        | 66cGTO    | 102.161    | 94.1592     | 97.8991     |

"Atomic units.

Table 2. The Calculated Potential Energy Function' of CIO+ with CI Energies

|                       | CISD/6-31G* | CISD/6-311G* | CISD/66cGTO |
|-----------------------|-------------|--------------|-------------|
| V <sub>0</sub> (a.u.) | -534.12750  | - 534.18627  | - 534.24742 |
| R, (Å)                | 1.48919     | 1.47771      | 1.46714     |
| $\int^2$              | 0.597458    | 0.548994     | 0.516785    |
| $f^3$                 | -1.88377    | - 1.83571    | -1.69073    |
| f4                    | 3.81835     | 3.980622     | 3.77578     |
| ſ                     | -21.2796    | - 19.2645    | - 18.8842   |
| ſ°                    | 102.111     | 80.2665      | 72.8912     |

a Atomic units.

of CIO molecule are well known, the same calculations of CIO molecule were carried out and the results were compared with experimental data. From these results of CIO it is possible to predict the range of errors in the values of various methods of calculations on CIO<sup>+</sup>.

**CIO.** For CIO radical both the experimental<sup>2-11</sup> and theoretical data<sup>21,22</sup> are available. The various levels of the approximation method were also applied for the calculations, using the basis sets of 6-31G<sup>\*</sup>, 6-311G<sup>\*</sup> and 66cGTO. The total energies of the MP and Cl calculations are shown in Figure 3 and 4, and their potential functions are in Table 4. The spectroscopic constants were calculated with same method used for ClO<sup>\*</sup>, and are shown in Table 5.

The bond lengths are longer than the experimental value for all methods of calculation. But, the bond length predicted by the basis set of 66cGTO is the closest to the experimental value among the various levels of calculation. In all cases the bond lengths were decreased about 0.02 Å with the basis set of 66cGTO, and consequently those values approach to the experimental value more closely. The SCF bond length shortens by 0.02Å upon correlation to MP3 and MP4DQ level.

For all methods using 6-31G<sup>\*</sup> basis set the vibration-rotation interaction constants,  $\alpha_e$ , are much smaller than the ex-

Table 3. The Calculated Spectroscopic Constants of CIO+



Figure 3. The potential curve of CIO with 6-311G\* basis set.



Figure 4. The potential curve of CIO with 66cGTO basis set.

| Method | Basis set | <i>V</i> <sub>0</sub> (a.u.) | $R_{r}(\text{\AA})$ | a,(MHz) | $D_t(KHz)$ | $B_{\ell}(\mathrm{GHz})$ | $\omega_{e}(\text{cm}^{-1})$ |
|--------|-----------|------------------------------|---------------------|---------|------------|--------------------------|------------------------------|
| SCF    | 6-31G*    | - 533.84676                  | 1.4841              | 295     | 38         | 20.90725                 | 1031.5                       |
|        | 6-311G*   | 533.88580                    | 1.4784              | 314     | 42         | 21.06710                 | 993.2                        |
|        | 66cGTO    | -533.91657                   | 1.4661              | 321     | 47         | 21.42283                 | 960.3                        |
| MP3    | 6-31G*    | - 534.13759                  | 1.4742              | 134     | 26         | 21.18996                 | 1264.3                       |
|        | 6-311G*   | -534.19723                   | 1.4644              | 161     | 31         | 21.47416                 | 1196.8                       |
|        | 66cGTO    | -534.26483                   | 1.4606              | 149     | 31         | 21.58651                 | 1204.9                       |
| MP4DQ  | 6-31G*    | -534.13927                   | 1.4738              | 140     | 27         | 21.19960                 | 1259.6                       |
|        | 6-311G*   | -534.19855                   | 1.4640              | 167     | 31         | 21.48596                 | 1193.1                       |
|        | 66cGTO    | - 534.26446                  | 1.4572              | 165     | 34         | 21.68541                 | 1155.2                       |
| MP4SDQ | 6-31G*    | - 534.14774                  | 1.4875              | 101     | 25         | 20.81072                 | 1267.8                       |
|        | 6-311G*   | -534.20823                   | 1.4776              | 119     | 28         | 21.09223                 | 1218.4                       |
|        | 66cGTO    | -534.27445                   | 1.4700              | 123     | 31         | 21.31074                 | 1178.2                       |
| CISD   | 6-31G*    | -534.12750                   | 1.4892              | 141     | 28         | 20.76416                 | 1199.4                       |
|        | 6-311G*   | -534.18627                   | 1.4777              | 164     | 32         | 21.08819                 | 1149.7                       |
|        | 66cGTO    | -534.24742                   | 1.4671              | 166     | 35         | 21.39299                 | 1115.5                       |

Spectroscopic Constants of ClO and ClO+

| Table 4 | . The | Potential | Energy | Function <sup>*</sup> | of | C10 |
|---------|-------|-----------|--------|-----------------------|----|-----|
|---------|-------|-----------|--------|-----------------------|----|-----|

|                       | Basis set | MP3         | MP4DQ       | MP4SDQ      | CISD        |
|-----------------------|-----------|-------------|-------------|-------------|-------------|
| V <sub>0</sub> (a.u.) | 6-31G*    | - 534.52477 | -534.52616  | - 534.53072 | - 534.51083 |
|                       | 6-311G*   | -534.58470  |             | - 534.59107 | - 534.57078 |
|                       | 66cGTO    | -534.65515  | - 534.65528 | - 534.66119 | 534.63360   |
| $R_{\epsilon}$ (Å)    | 6-31G*    | 1.60842     | 1.60882     | 1.61215     | 1.61479     |
|                       | 6-311G*   | 1.60725     | 1.60782     | 1.61056     | 1.61204     |
|                       | 66cGTO    | 1.58623     | 1.58506     | 1.58698     | 1.59015     |
| $f^2$                 | 6-31G*    | 0.347779    | 0.338737    | 0.334696    | 0.292768    |
|                       | 6-311G*   | 0.246195    | 0.242699    | 0.242827    | 0.263489    |
|                       | 66cGTO    | 0.290437    | 0.287197    | 0.290226    | 0.291553    |
| $f^3$                 | 6-31G*    | -0.643851   | -0.654212   | -0.607065   | -0.712936   |
|                       | 6-311G*   | -0.743493   | -0.735135   | - 0.743564  | -0.768710   |
|                       | 66cGTO    | -0.883791   | -0.881686   | -0.909525   | -0.813813   |
| $f^4$                 | 6-31G*    | 0.066199    | 0.069009    | 0.472128    | 0.840000    |
| -                     | 6-311G*   | 1.20219     | 1.27600     | 1.35496     | 0.621298    |
|                       | 66cGTO    | 2.65744     | 2.71932     | 2.77999     | 2.74234     |
| $f^5$                 | 6-31G*    | - 13.1466   | - 12.7572   | -13.8427    | -10.4063    |
| -                     | 6-311G*   | -8.55671    | - 8.65944   | -8.58039    | -6.84741    |
|                       | 66cGTO    | - 11.0389   | -11.0847    | - 10.6135   | -11.6658    |
| ſ°                    | 6-31G*    | 46.4210     | 46.0008     | 35.3628     | 42.4243     |
|                       | 6-311G*   | 39.2387     | 38.1906     | 36.5317     | 55.9422     |
|                       | 66cGTO    | 26.2488     | 25.6750     | 23.5201     | 25.7362     |

a Atomic units.

Table 5. The Calculated Spectroscopic Constants of CIO

| Method | Basis set | V <sub>0</sub> (a.u.) | <i>R</i> , (Å) | a, (MHz) | D <sub>e</sub> (KHz) | B, (GHz) | $\omega_r$ (cm <sup>-1</sup> ) |
|--------|-----------|-----------------------|----------------|----------|----------------------|----------|--------------------------------|
| SCF    | 6-31G*    | - 534.23176           | 1.6220         | 307      | 44                   | 17.50277 | 738                            |
|        | 6-311G*   | - 534.27301           | 1.6310         | 48       | 34                   | 17.31070 | 826                            |
|        | 66cGTO    | - 534.29668           | 1.6083         | 67       | 33                   | 17.80199 | 863                            |
| MP3    | 6-31G*    | -534.52477            | 1.6084         | 61       | 30                   | 17.79988 | 915                            |
|        | 6-311G*   | -534.58470            | 1.6072         | 170      | 43                   | 17.82581 | 770                            |
|        | 66cGTO    | -534.65515            | 1.5862         | 164      | 39                   | 18.30143 | 836                            |
| MP4DQ  | 6-31G*    | -534.52616            | 1.6088         | 67       | 31                   | 17.79096 | 903                            |
|        | 6-311G*   | - 534.58594           | 1.6078         | 172      | 43                   | 17.81323 | 764                            |
|        | 66cGTO    | 534.65528             | 1.5851         | 167      | 40                   | 18.32841 | 832                            |
| MP4SDQ | 6-31G*    | -534.53072            | 1.6121         | 59       | 31                   | 17.71767 | 898                            |
|        | 6-311G*   | -534.59107            | 1.6106         | 174      | 43                   | 17.75263 | 765                            |
|        | 66cGTO    | -534.66119            | 1.5869         | 171      | 39                   | 18.28407 | 836                            |
| CISD   | 6-31G*    | -534.51083            | 1.6148         | 110      | 35                   | 17.65973 | 840                            |
|        | 6-311G*   | - 534.57078           | 1.6120         | 155      | 39                   | 17.72001 | 796                            |
|        | 66cGTO    | -534.66360            | 1.5901         | 142      | 38                   | 18.21125 | 838                            |
| exp."  |           |                       | 1.570          | 174      | 39                   | 18.60216 | 854                            |

<sup>e</sup>References 9 and 32.

perimental value while the MP3, MP4DQ and MP4SDQ values using 6-311G<sup>\*</sup> and 66cGTO basis sets are close to the experimental value. The centrifugal distortion constants,  $D_{e}$ , are all in reasonable agreement with the experimental values. The harmonic frequencies of the MP3, MP4DQ, and MP4SDQ with 6-31G<sup>\*</sup> are overestimated by about 5% and underestimated by the same amount with 6-311G<sup>\*</sup> basis set, but estimated correctly by all methods with 66cGTO. The CI harmonic frequencies approach closely to experimental

values.

## Discussion

Various potential energy functions have been calculated by different orders of M $\phi$ ller-Plesset perturbation theory and by the CI method for ClO and ClO<sup>+</sup>. From the molecular parameters of ClO, the reasonable agreement with experimental value was found in the calculations with 66cGTO 480 Bull. Korean Chem. Soc., Vol. 14, No. 4, 1993

basis set. In bond lengths the deviation of less than 1% is good agreement. Specially the MP3 and MP4DQ calculations with the 66 cGTO basis set gave good agreements over all spectroscopic parameters. The values from CI calculations are no better than the MP values, which is different result from our previous calculations. But the MP3 calculation gave reasonably good results as usual. The previous theoretical values<sup>22</sup> also showed longer bond lengths. But the values of  $\alpha_{e}$ , and  $D_{e}$  gave good agreement with experimental values.

In all MP levels of calculation the bigger basis set gave better results, which was expected from the previous calculations<sup>30</sup>.

Therefore the spectroscopic parameters of CIO<sup>+</sup> calculated from the method including electron correlation with large basis set will give the least discrepancies from the experimental values.

Acknowledgement. This work was supported by NON DIRECTED FUND, Korea Research Foundation, 1990.

#### References

- M. Trolier, R. L. Mauldin III, and A. R. Ravishanakra, J. Phys. Chem., 94, 4896 (1990).
- 2. G. Pannetier and A. G. Gaydon, Nature, 161, 242 (1948).
- 3. G. Porter, Discuss. Faraday Soc., 9, 60 (1950).
- 4. R. A. Durie and D. A. Ramsay, Can. J. Phys., 36, 35 (1958).
- 5. J. A. Coxon and D. A. Ramsay, *Can, J. Phys.*, 54, 1034 (1976).
  - A. Loewenschuss, J. C. Millar, and L. Andrews, J. Mol. Spectrosc., 81, 351 (1980).
  - M. T. Duignan and J. W. Hudgens, J. Chem. Phys., 82, 4426 (1985).
  - 8. T. Amano and E. Hirota, J. Mol. Spectrosc., 66, 185 (1977).
  - R. K. Kakar, E. A. Cohen, and M. Geller, J. Mol. Spectrosc., 70, 243 (1978).
  - J. B. Burkholder, P. D. Hammer, C. J. Howard, A. G. Maki, G. Thompson, and C. Chackerian, Jr., J. Mol. Spectrosc., 124, 139 (1987).
  - V. I. Lang, S. P. Sander, and R. R. Friedl, J. Mol. Spectrosc., 132, 89 (1988).
  - 12. E. M. Weinstock, M. J. Phillips, and J. G. Anderson, J.

Hae-Sun Song and Eun-Mo Sung

Geophys, Res., 86, 7273 (1981).

- W. H. Brune, E. M. Weinstock, and J. G. Anderson, *Geophys. Res. Lett.*, 15, 144 (1988).
- P. M. Solomon, R. L. deZafra, A. Parrish, and J. W. Barrett, *Science*, 224, 1210 (1984).
- B. Carli, F. Mencaraglia, M. Carlotti, B. M. Dinelli, and I. Nolt, J. Geophys. Res., 93, 7063 (1988).
- R. L. deZafra, M. Jaramillo, J. Barrett, L. K. Emmons, P. M. Solomon, and A. Parrish, *J. Geophys. Res.*, 94, 11423 (1989).
- 17. R. T. Menzies, Geophys, Res., Lett., 10, 729 (1983).
- T. Kostiuk, J. L. Faris, M. J. Mumma, D. Deming, and J. J. Hillman, J. Geophys. Res., 91, 2735 (1986).
- 19. O. F. Raper, C. B. Faramer, R. Zander, and J. H. Park, J. Geophys. Res., 92(DA), 9851 (1987).
- P. A. G. O'Hare and A. C. Wahl, J. Chem. Phys., 54, 3770 (1971).
- 21. J. Lievin and J. Metz, Theoret. Chim. Acta, 67, 391 (1985).
- L. G. M. Pettersson, S. R. Langhoff, and D. P. Chong, J. Chem. Phys., 85, 2836 (1986).
- V. L. Pershin, A. I. Boldyrev, and O. P. Charkin, Zh. Fiz. Khim., 61, 120 (1987).
- 24. T. Amano, J. Chem. Phys., 79, 3595 (1983).
- 25. B. E. Turner, Astrophys. J., 376, 573 (1991).
- P. F. Bernath, S. A. Rogers, L. C. O'Brien, C. R. Brazier, and A. D. McLean, *Phys. Rev. Lett.*, **60**, 197 (1988).
- S. R. Langhoff and C. W. Bauschlicher, Jr., Astrophys. J., 375, 843 (1991).
- M. A. Castro, S. Canuto, and F. M. Plathe, Astrophts. J., 367, L69 (1991).
- W. J. Hehre, L. Radom, P. V. R. Schleyer, and J. A. Pople, "Ab initio Molecular Orbital Theory", John Wiley & Sons, Inc. New York, 1986.
- R. Krishnan, M. J. Frish, and J. A. Pople, *J. Chem. Phys.*, 72, 4244 (1980).
- 31. E. M. Sung, Bull. Korean. Chem. Soc., 10, 377 (1989).
- J. A. Coxon, W. E. Jones, and E. G. Skolnik, *Can. J. Phys.*, 54, 1043 (1976).
- 33. S. Huzinaga, J. Chem. Phys., 42, 1293 (1965).
- 34. A. Veillard, Theor, Chim, Acta., 12, 405 (1968).
- 35. P. Rosmus and W. Meyer, J. Chem. Phys., 63, 2356 (1975).