

in total cloud complex mass and mass supply rate, when these physical quantities are confined by observed zodiacal light brightness and dust particle SFD at 1 au. We noticed that, if we assume the existence of fluffy aggregates discovered in the Earth's stratosphere and the coma of 67P/Churyumov-Gerasimenko, the required mass supply rate decreases significantly. We also found out that close encounters with planets (mostly Jupiter) are the dominating factor of the orbital evolution of dust particles, as the result, the lifetime of cometary dust particles are shorter than Poynting-Robertson lifetime (around 250 thousand years). As another consequence of severe close encounters, only a small fraction of cometary dust particles can be transferred into the orbit  $< 1$  au. This effect is significant for large size particles of  $\beta < 0.01$ . The exceptional cases are dust particles ejected from 2P/Encke and active asteroids. Because they rarely encounter with Jupiter, most dust particles ejected from those objects are governed by Poynting-Robertson effect and well transferred into the orbits of small semimajor axis. In consideration of the above effects, we directly estimated probability of mutual collisions between dust particles and concluded that mutual collisions in the IDPs cloud complex is mostly ignorable, except for the case of large sized particles from active asteroids.

#### [구 SO-02] Spectropolarimetric Signals of Comet 2P/Encke During Its 2017 Apparition

Yuna Grace Kwon<sup>1</sup>, Masateru Ishiguro<sup>1</sup>, Daisuke Kuroda<sup>2</sup>, Koji S. Kawabata<sup>3</sup>, Tatsuya Nakaoka<sup>3</sup>, Miho Kawabata<sup>3</sup>, Makoto Uemura<sup>3</sup>, Yoshiharu Shinnaka<sup>4</sup>, and Hiroshi Akitaya<sup>5</sup> and OISTER collaboration team

<sup>1</sup>Seoul National University, Korea;

<sup>2</sup>ynkwon@astro.snu.ac.kr, <sup>3</sup>Okayama Astrophysical Observatory, Japan, <sup>4</sup>Hiroshima University, Japan,

<sup>5</sup>National Astronomical Observatory of Japan,

<sup>5</sup>Center for Astronomy, Ibaraki University, Japan

Comets are one of the most primordial solar system objects that hold the information of the early days of solar system formation inside their nuclei. Orbiting the Sun, they spew such ancient materials that have been buried for many years, creating dust and gas comae. Cometary dust grains absorb and scatter sunlight radiating the continuous light, while gas molecules form the line emissions. Each of the comets has its own light patterns, which depends on the physical and chemical properties of the dust and gas components. In this regard, spectropolarimetry can

be a powerful tool to study the properties of cometary constituents free from contamination of each other. This methodology offers a series of information on the polarization degrees of the dust and gas components as well as on wavelength dependence of the polarization degree and polarization angle of cometary dust simultaneously. Herein, we will report the results of the spectropolarimetric study of comet 2P/Encke, which is one of the well-known objects for its shortest orbital period and its prominent aging signals. We performed a spectropolarimetric observation of comet 2P/Encke in its inbound orbit using the Higashi-Hiroshima Optical and Near-Infrared Camera (HONIR) at the Higashi-Hiroshima Observatory, Japan, on UT 2017 February 21 at high phase angle of  $\approx 75.7$  deg. Our study of this interesting comet is the first and only one done through spectropolarimetry in a referred publication. We will discuss the most recent polarimetric results of our study in terms of 2P/Encke's current evolutionary status.

#### [구 SO-03] The fragmented asteroid 354P/LINEAR (2010 A2) captured by the K-GMT science program

Yoonyoung Kim, Masateru Ishiguro, Myung Gyoon Lee

Seoul National University

With support from the K-GMT science program (PID: GN-2016B-Q-14), we conducted observations of active asteroid 354P/LINEAR (2010 A2) when it made its closest approach to Earth (i.e., the geocentric distance of 1.06 au on 2017 January 27-28). Taking advantage of the best observing geometry since the discovery, we obtained the first evidence for the rotational status of the largest fragment (~120 m in diameter), which was slowly rotating, that is, the rotational period of 11.36 hours. In addition, we succeed in direct imaging of 10 sub-fragments (~20 m in diameter or larger). Based on these new observational results, we conjecture that this active asteroid was created as a result of catastrophic collision among unknown asteroids. The details of this work are given in *Astrophysical Journal Letters*, 842, L23.

#### [구 SO-04] Thermal Modeling of Comet-Like Objects from AKARI Observations

Yoonsoo P. Bach<sup>1</sup>, Masateru Ishiguro<sup>1</sup>, Fumihiko Usui<sup>2</sup>

<sup>1</sup>Department of Physics and Astronomy, Seoul