

# Age Distribution of Galactic Globular Clusters using HST Snapshot photometry

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We present relative ages for a sample of Galactic globular clusters (GCs) using their color-magnitude diagrams (CMDs) observed with the HST/WFPC2 camera in the F439W and F555W bands. The ages have been obtained by a differential comparison of the CMDs of GCs using  $\Delta(B-V)$  method, the color difference between main-sequence turnoff and the lower red-giant branch. All metal-poor GCs with  $[Fe/H] < -1.7$  show old ( $\sim 12$  Gyr) ages and are coeval. All the metal-rich GCs with  $[Fe/H] > -0.8$  are found to be  $\sim 0.8$  Gyr younger than the most metal-poor ones. Intermediate-metallicity clusters ( $-1.7 < [Fe/H] < -0.8$ ) are on average 2 Gyr younger than the most metal-poor counterparts, with a large age dispersion and a total age range of  $\sim 2$  Gyr. We also discuss the correlation of relative ages with the horizontal-branch morphologies of GCs.