## Observations of Molecular Hydrogen in the Carina Nebula

Dae-Hee Lee<sup>1</sup>, Mark Hurwitz<sup>2</sup>, Kyoung-Wook Min<sup>1,3</sup>

<sup>1</sup> Department of Physics, Korea Advanced Institute of Science and Technology,
373-1, Kusong-dong, Yusong-gu, Taejon 305-701, Korea

<sup>2</sup> Space Sciences Laboratory, University of California, Berkeley, CA94720-5030, USA

<sup>3</sup> Satellite Technology Research Center, Korea Advanced Institute of Science and
Technology, 373-1, Kusong-dong, Yusong-gu, Taejon 305-701, Korea

Observations of molecular hydrogen absorption lines in the continuum spectra of 3 early-type stars in the Carina Nebula, HD93129A, HD93250, and HD303308, were made with the Berkeley EUV/FUV spectrometer on the ORFEUS telescope in 1993 September. Using high-resolution optical observations of Na I absorption lines, with the constraint on the distribution and velocity of molecular clouds along each line of sight, we obtain column densities for each  $H_2$  rotational level and derive excitation temperatures and UV radiation fields for the  $H_2$  clouds towards each star. All three stars show strong  $H_2$  absorption features, indicating that the Carina Nebula contains abundant hydrogen molecules. Also, the UV radiation field strength is found to be correlated with the distance between the target star and  $\eta$  Car. Based on these results, we examine the morphology of the Carina Nebula in association with the CO molecular clouds.